



國立清華大學
NATIONAL TSING HUA UNIVERSITY

NCTS

PRPC
PHYSICS RESEARCH PROMOTION CENTER

Basic hydrodynamics and Introduction to FLASH

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Outline:

- Astrophysical fluids
- Basic hydrodynamic equations
- Introduction to the FLASH code



Why we need simulations in Astronomy?

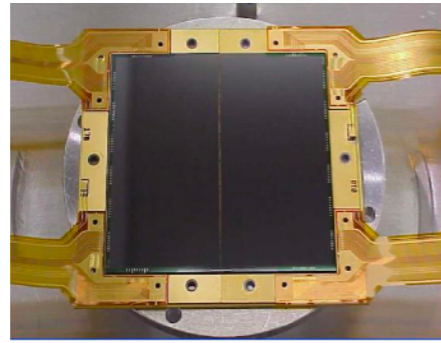
- We **receive** and **record**, but we do not experiment or perturb (too distant)
- Astrophysical systems involve extreme conditions that cannot be replicated in lab (mostly)
- Timescales are \gg human lifetime (mostly)



Observation and Computation Analogies



Astrophysical object



Detectors

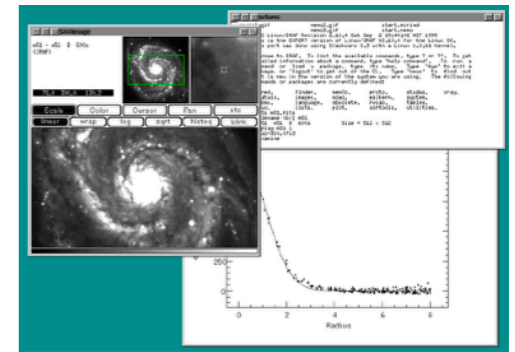


Telescope

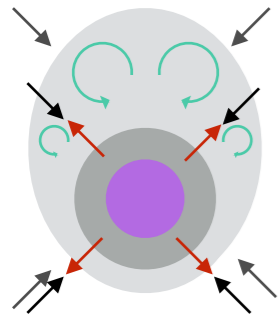
Observation



DATA



Data reduction / Calibration



Mathematical Model

```
subroutine Mod (S10m, S10m, dt, time)
  use Inflow, Outflow, Diff, Pressure, Gravity, Opacity, Cooling, Heating
  use Driver_Interface, ONLY: Driver_Init, Driver_TimeStep,
  use Grid_Interface, ONLY: Grid_Init, Grid_GetNeighbors, Grid_GetType, Grid_GetVolume,
  use TagFile_Interface, ONLY: TagFile_Write
  implicit none
  integer, intent(in) :: S10m
  integer, intent(in) :: dt
  real, intent(in) :: time
  real, dimension(S10m) :: S10m
  real, dimension(S10m) :: dt
  real, dimension(S10m) :: time
  ! ... (rest of the code) ...
end subroutine Mod
```

Codes

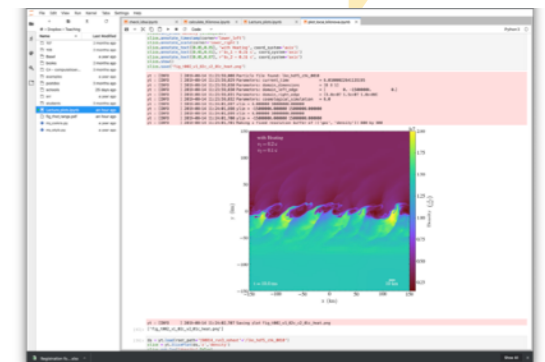


Supercomputer

Simulation



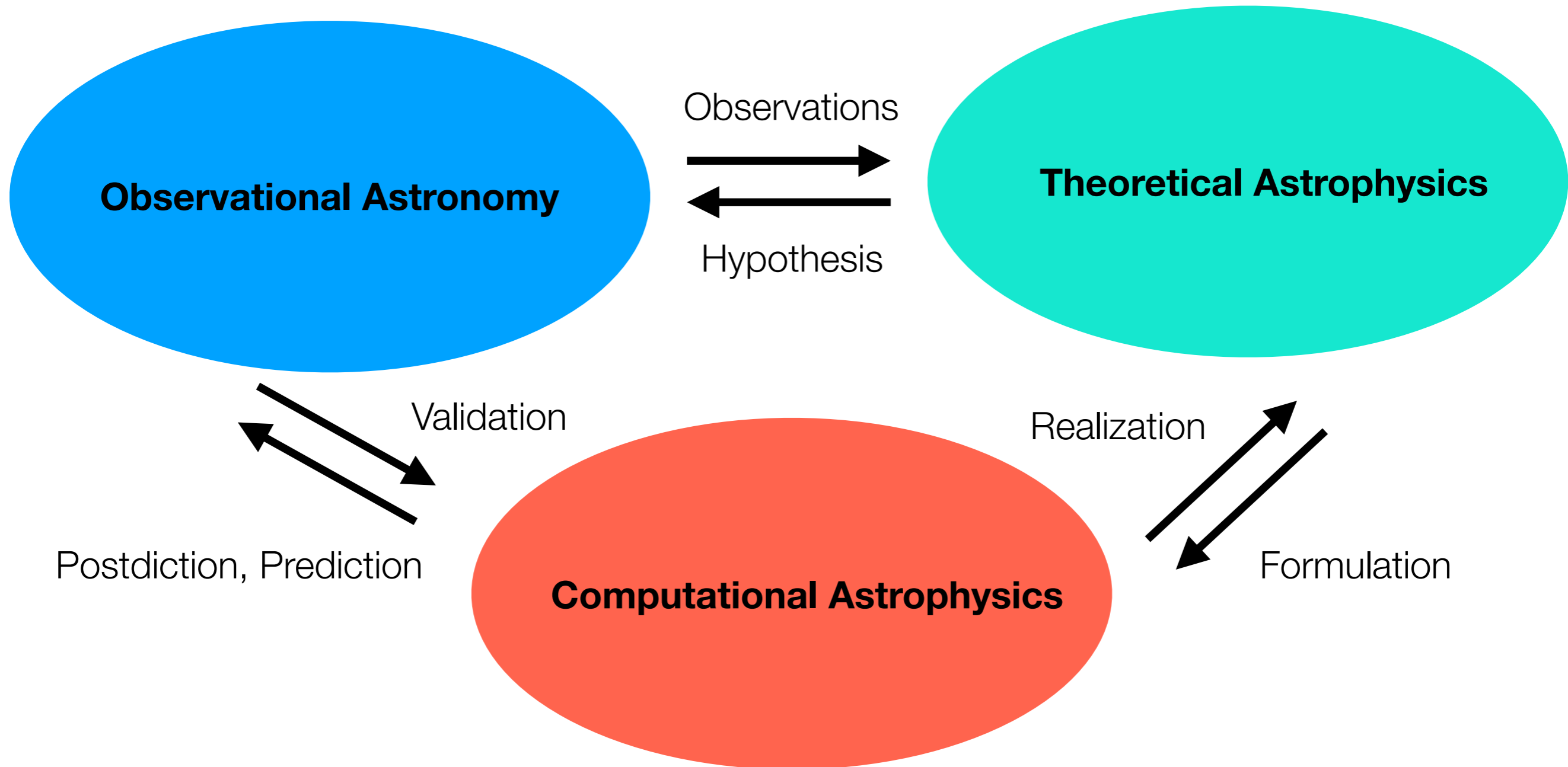
DATA



Data analysis / Visualization



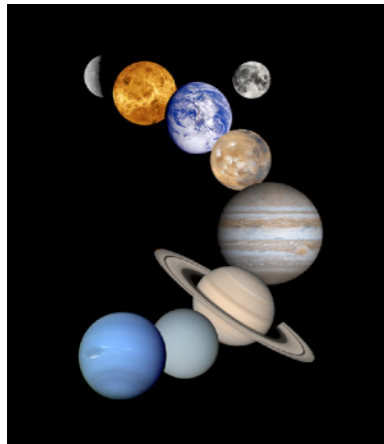
Frontier Astronomy Research



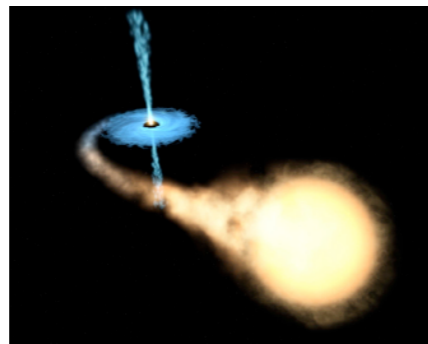
Michael Norman (1997)



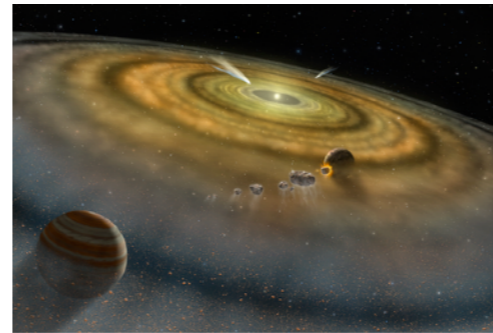
Astrophysical Problems



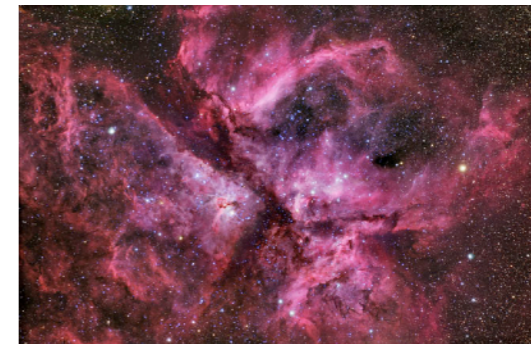
Planets
(~ Earth radius)



Stars/Binaries
(~0.1-1000 AU)



Planetary disk
(~100-1000 AU)



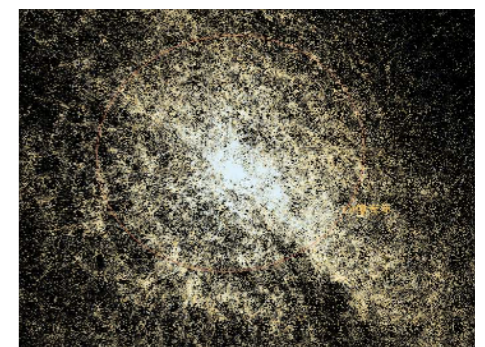
Interstellar medium
(~10-100 pc)



Galaxy
(~50 kpc)



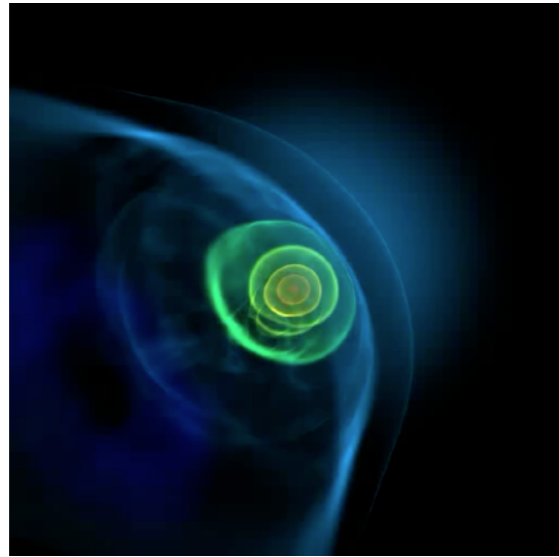
Galaxy clusters
(~1 Mpc)



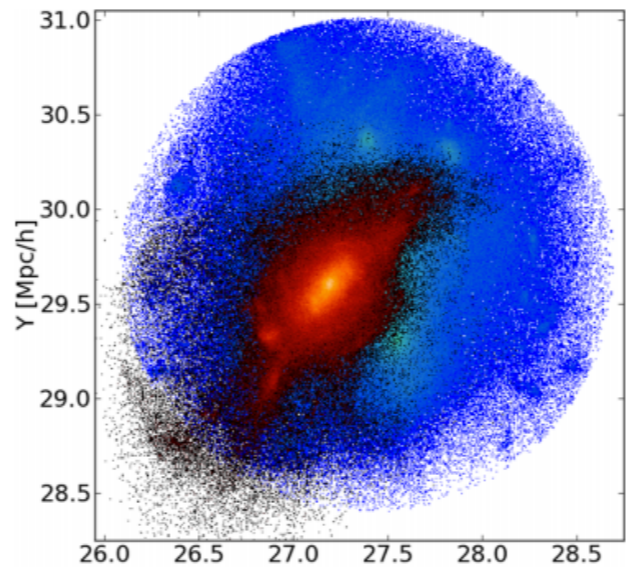
Large scale structure
(~1 Gpc)



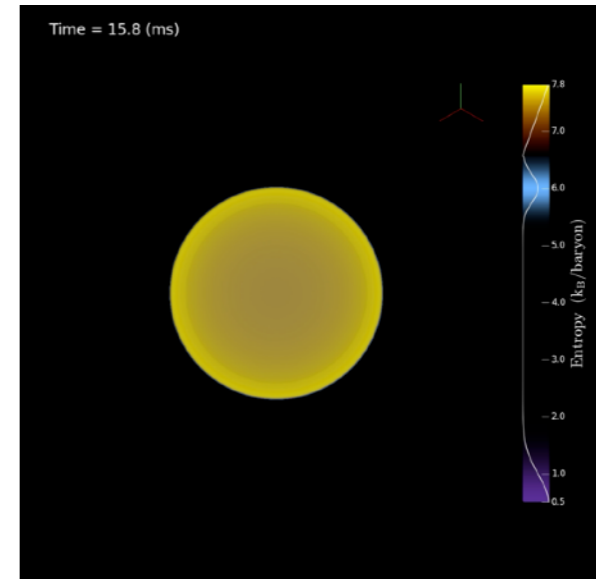
Astrophysical Simulations



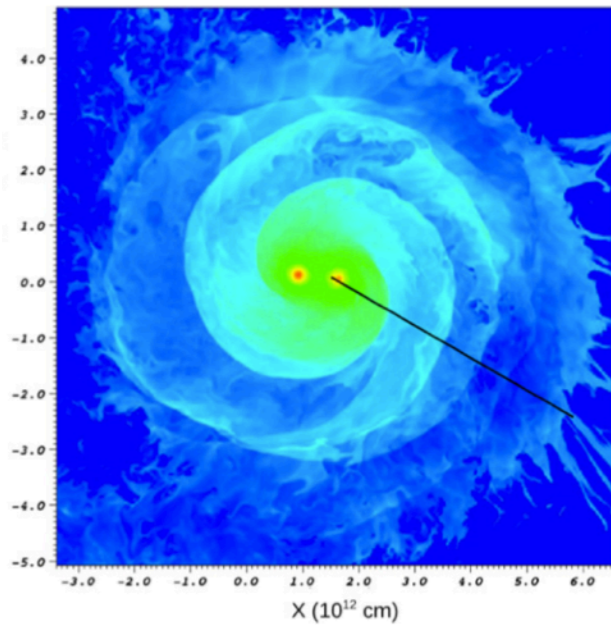
2012 ApJ, 750, 151



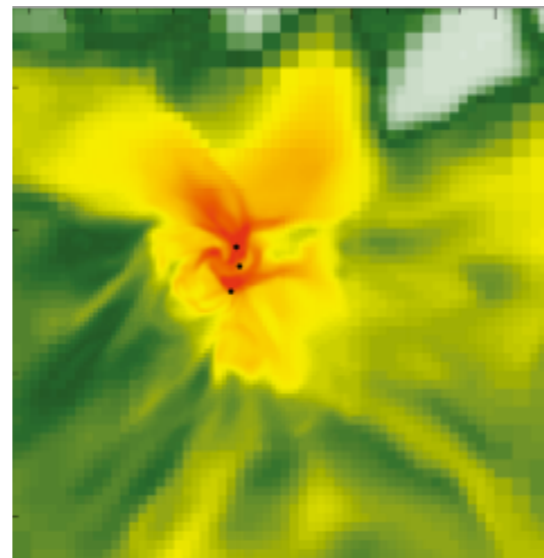
2013MNRAS



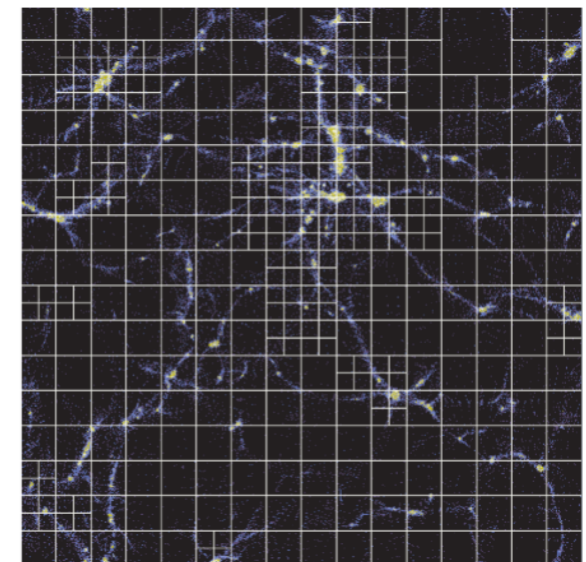
2018 ApJ



2012 ApJ, 746, 74



2009 MNRAS 398, 1082



2005 ApJS 160, 28



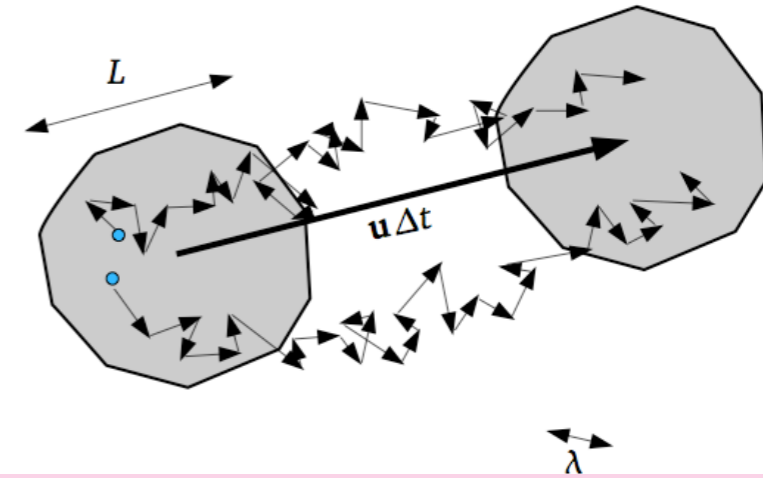
Fluid Approximation

- To solve the gas dynamics equations, one could use **N-body** method if we know the force between two particles.
- However, the number of particles is usually huge ($N_A=10^{23}$). It is impossible to use direct N-body method to study the gas dynamics. **-> Fluid**

Fluid Approximation (continue)

$L \gg \lambda_{\text{mfp}}$, Mean free path

$T \gg \tau_{\text{coll}}$, Collision time scale



Exercise: Consider the air in this room

$$\rho \sim 10^{-3} \text{ g/cm}^3, \mu \sim 28m_p, T \sim 300 \text{ K}$$

$$\text{Cross section } \sigma \sim 3 \times 10^{-15} \text{ cm}^2$$

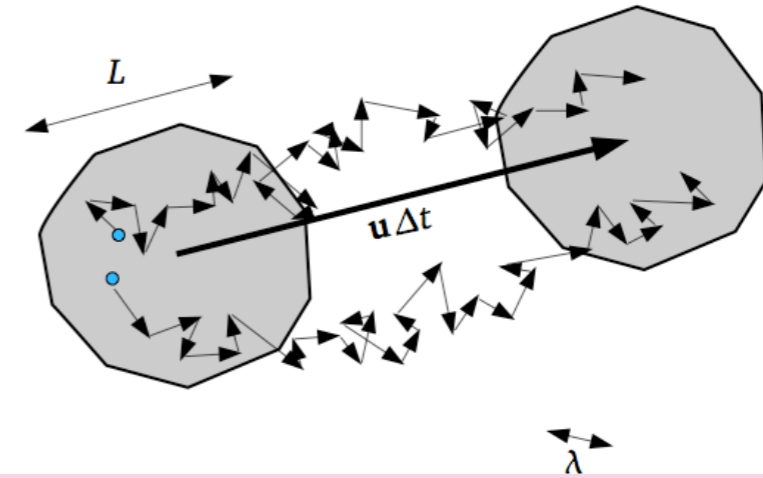
$$\text{RMS speed of nitrogen } v \sim \sqrt{\frac{8kT}{\pi\mu}} \sim 5 \times 10^4 \text{ cm/s}$$

$$\tau_{\text{coll}} \sim (n\sigma v)^{-1} \sim 3 \times 10^{-10} \text{ sec} \quad \lambda_{\text{mfp}} \sim (n\sigma)^{-1} \sim 3 \times 10^{-5} \text{ cm}$$

Fluid Approximation (continue)

$L \gg \lambda_{\text{mfp}}$, Mean free path

$T \gg \tau_{\text{coll}}$, Collision time scale



Exercise: In ISM

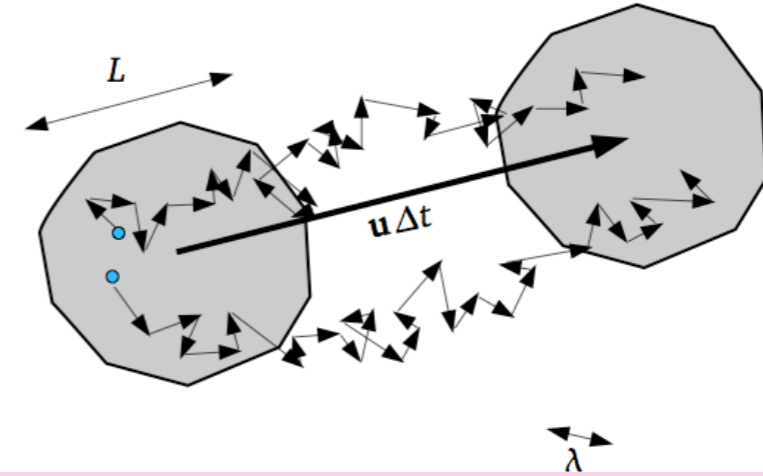
Cross section $\sigma \sim ? \text{ cm}^2$

number density $n \sim ? \text{ cm}^{-3}$

Fluid Approximation (continue)

$L \gg \lambda_{\text{mfp}}$, Mean free path

$T \gg \tau_{\text{coll}}$, Collision time scale



Exercise: In ISM

Cross section $\sigma \sim 3 \times 10^{-15} \text{ cm}^2$

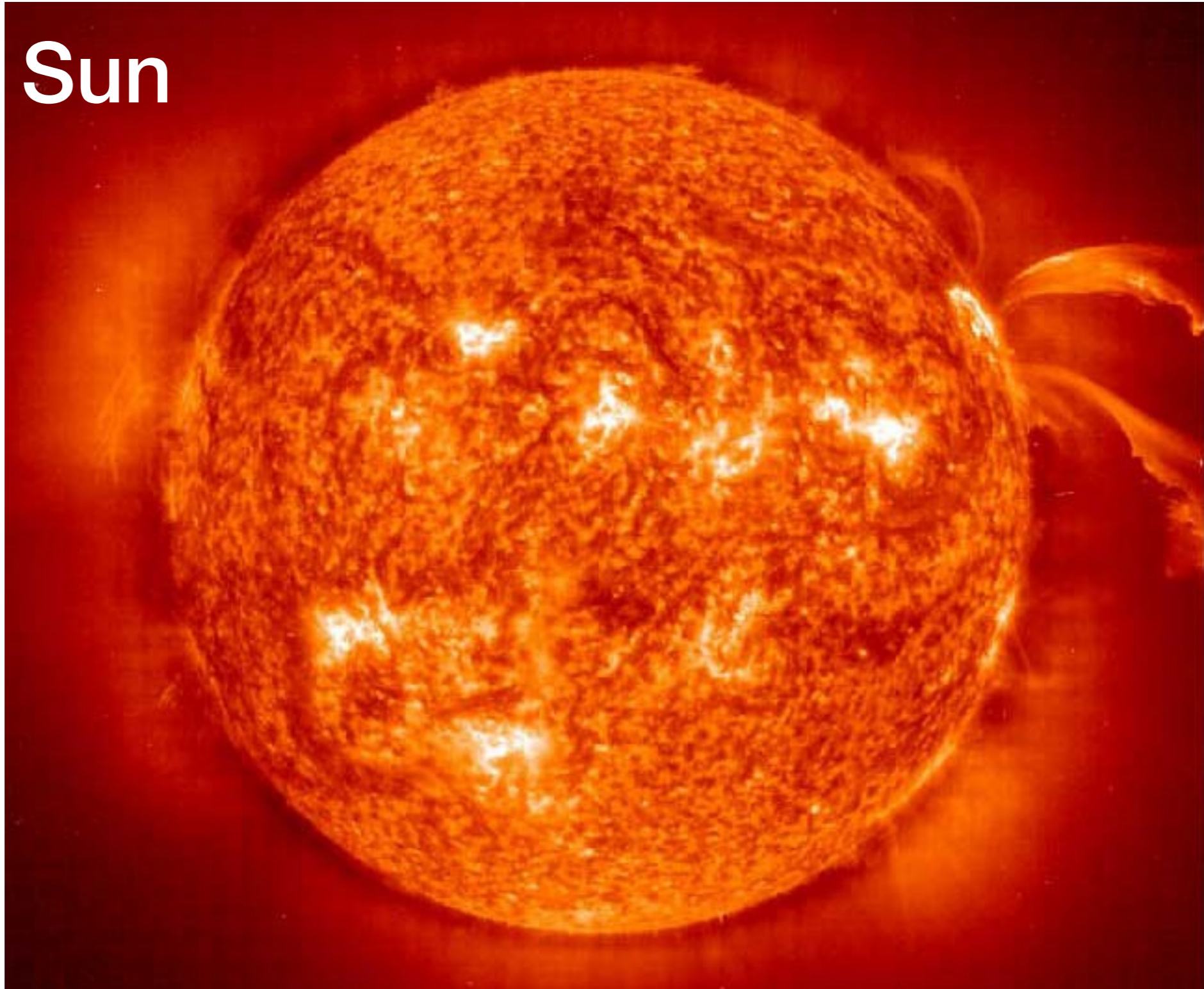
number density $n \sim 1 \text{ cm}^{-3}$

$\lambda_{\text{mfp}} \sim (n\sigma)^{-1} \sim 10^{15} \text{ cm}$

$\tau_{\text{coll}} \sim (n\sigma v)^{-1} \sim 10^9 \text{ sec} \sim 40 \text{ yrs}$



Sun





Star forming region (S106)

Image credit: NASA

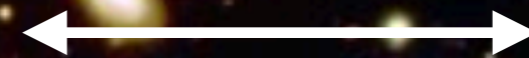
Cas A Supernova Remnant

Image credit: NASA

Bullet cluster

Image credit: NASA

0.5 Mpc

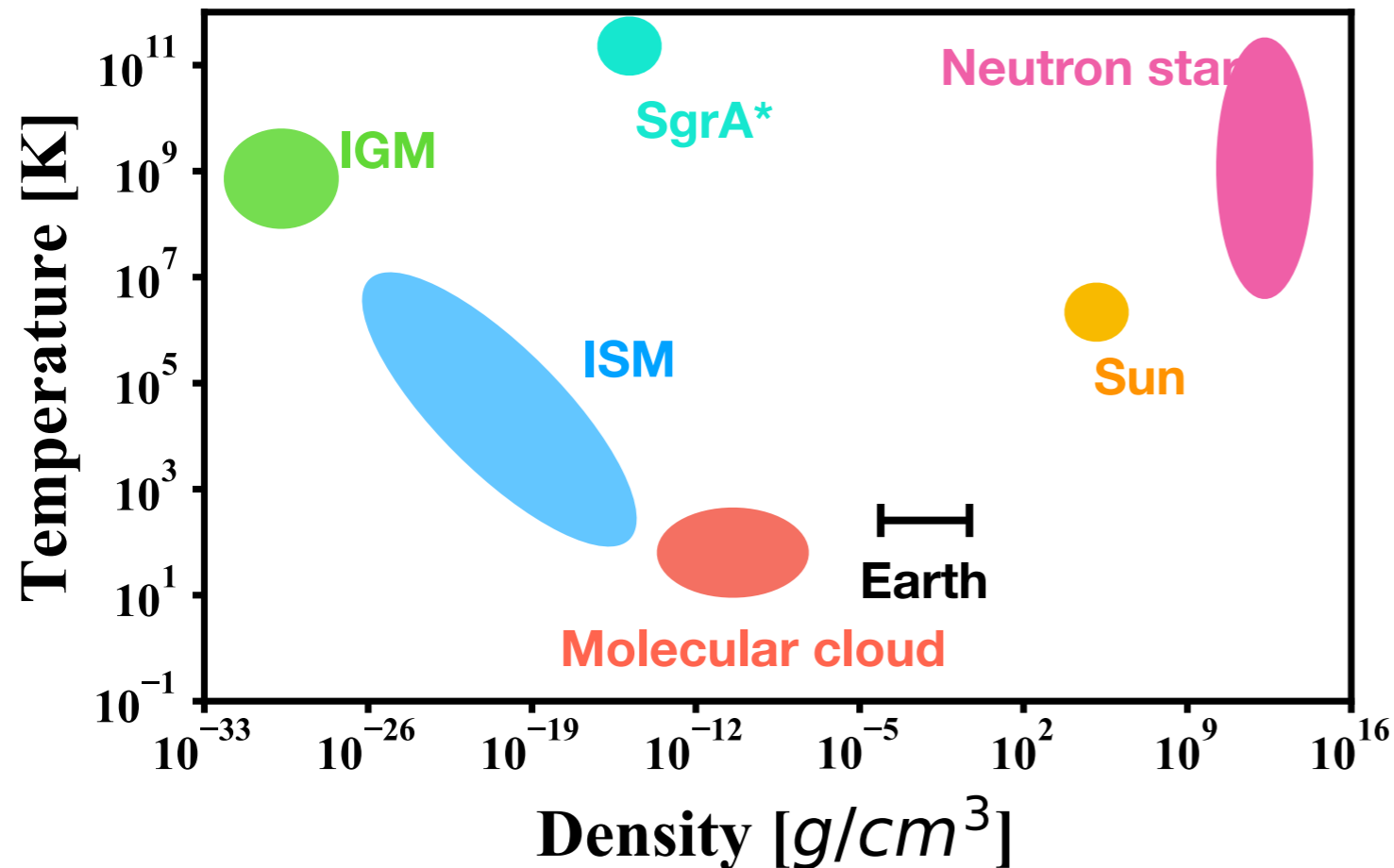


Astrophysical Fluids



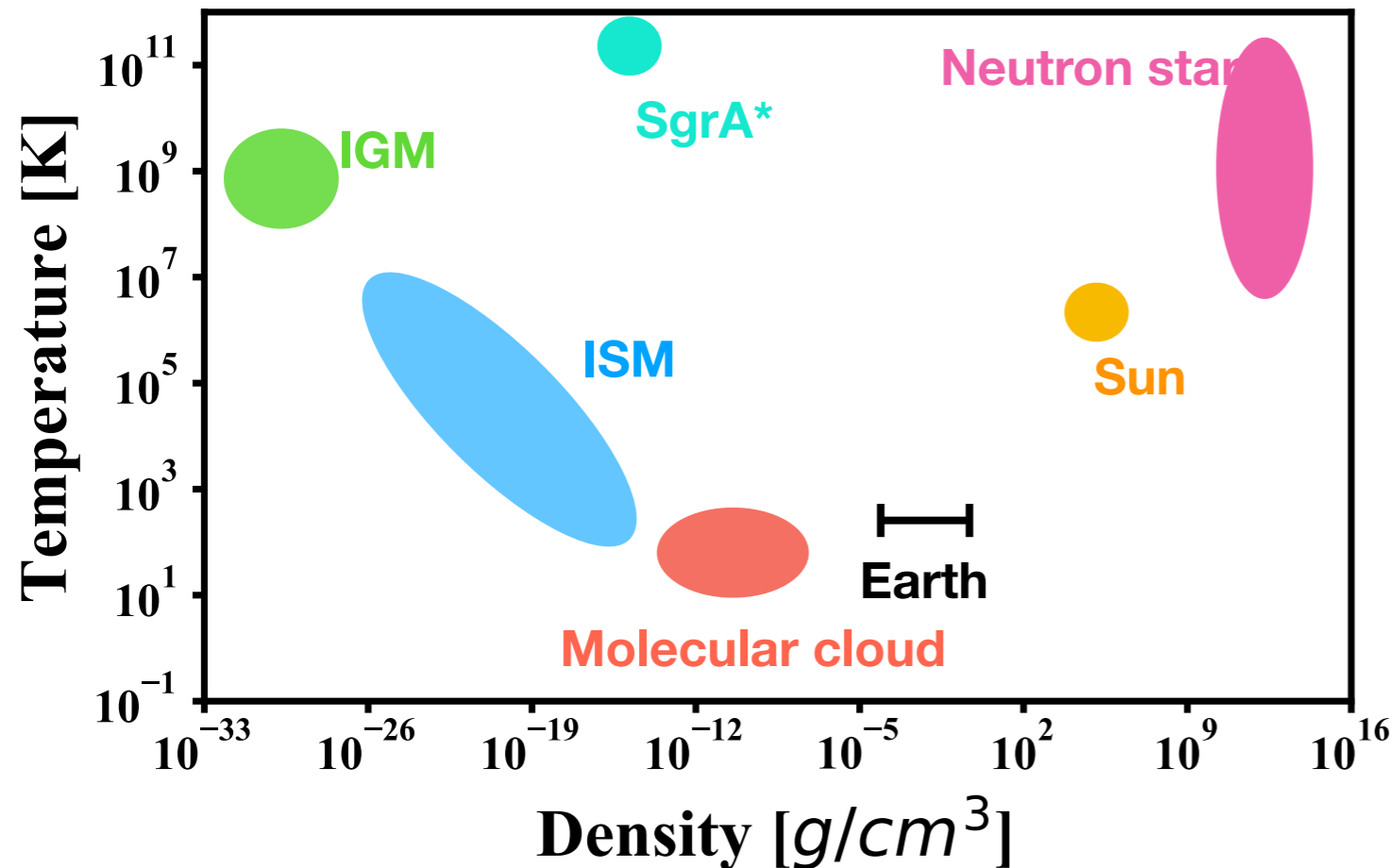
<https://youtu.be/ICxC5ekWnUc>

Astrophysical fluids



- Different with Earth-bound fluid
- Wide range of physical Conditions

Astrophysical fluids



- **Magnetic fields** are often important
- **Compressibility** is often important
- Occasionally **self-gravitating** and **relativistic**
- Non-ideal processes are often unimportant
- Sometimes involves interactions with non-thermal processes and strong radiation fields



Governing equations

The governing equations for ideal (inviscid) hydrodynamics. The main physical ideas are simple:

- **Mass** is conserved (1 constraint)
- **Momentum** is conserved (3 constraints)
- **Energy** is conserved (1 constraint)

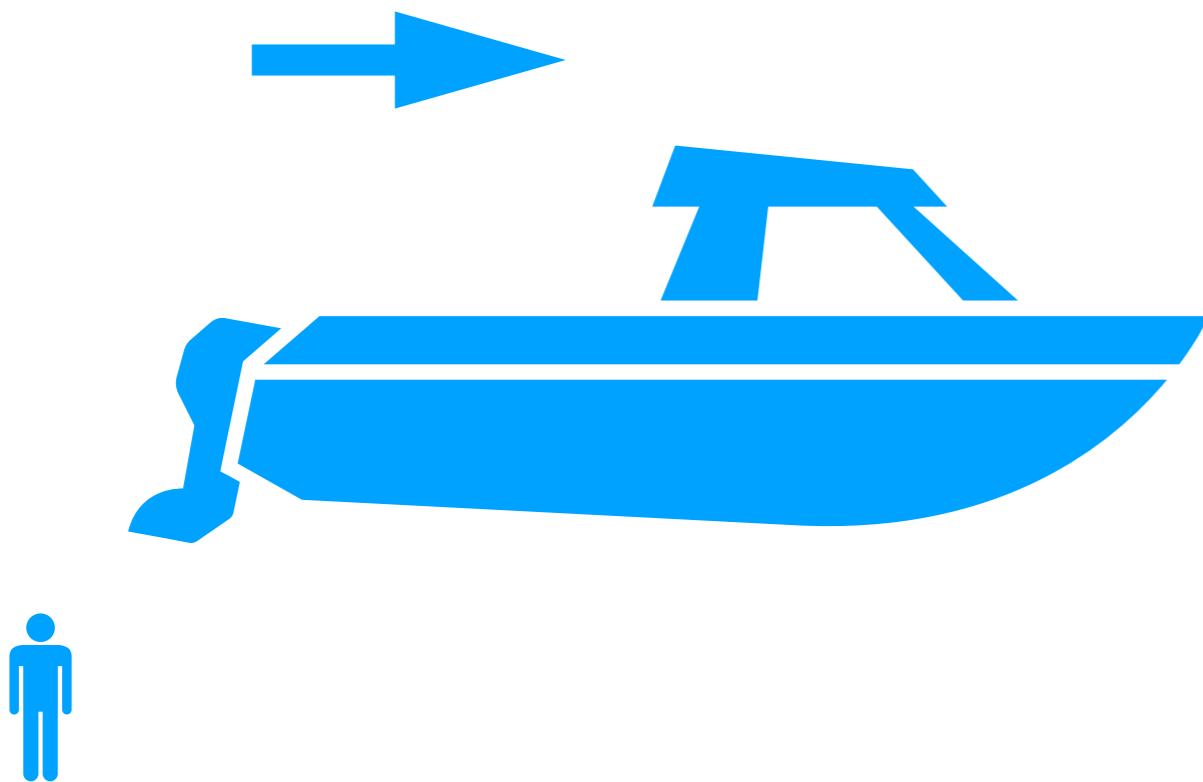
variables: ρ , \mathbf{v} , and u

ρ	Gas density
\mathbf{v}	Gas velocity
u	Gas internal energy



Coordinates

Eulerian



Lagrangian



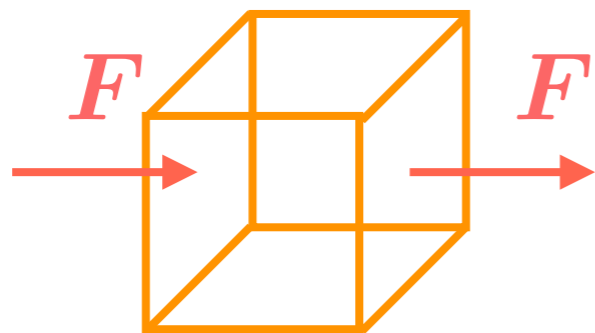
Conservational laws

The conservational laws can be written in what is called “**conservation form**”:

$$\partial_t(U) = -\nabla \cdot \mathbf{F}$$

“Gauss’s Theorem” or “divergence theorem”

$$\int_V (\nabla \cdot \mathbf{F}) dV = \oint_S \mathbf{F} \cdot \mathbf{n} dS$$



U Density of a quantity
 \mathbf{F} Flux density for that quantity



Continuity equation

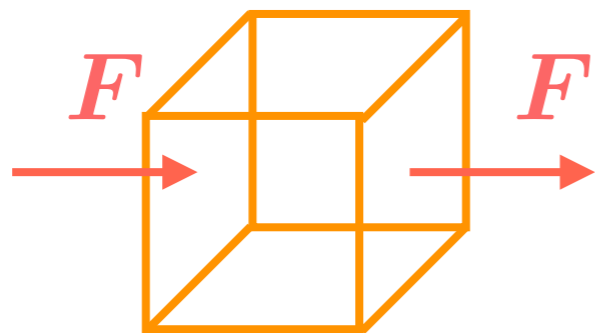
$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

ρ Gas density

\mathbf{v} Gas velocity

"Gauss's Theorem" or "divergence theorem"

$$\int_V (\nabla \cdot \mathbf{F}) dV = \oint_S \mathbf{F} \cdot \mathbf{n} dS$$





Convective derivative

- Consider a fluid element

$$\mathbf{v} = \frac{d\mathbf{r}}{dt}, \quad \mathbf{a} = \frac{d\mathbf{v}}{dt},$$

$$\begin{aligned} \mathbf{a} &= \frac{d\mathbf{v}(\mathbf{r}, t)}{dt} \\ &= \frac{\partial \mathbf{v}}{\partial t} + \frac{\partial \mathbf{v}}{\partial r} \cdot \frac{\partial \mathbf{r}}{\partial t} \\ &= \frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} \\ &\equiv \frac{D\mathbf{v}}{Dt} \end{aligned}$$

Convective derivate

- Rewrite continuity equation

$$\begin{aligned} \partial_t \rho &= -\nabla \cdot (\rho \mathbf{v}) \\ &= -(\mathbf{v} \cdot \nabla) \rho - \rho(\nabla \cdot \mathbf{v}) \end{aligned}$$

$$\rightarrow \partial_t \rho + (\mathbf{v} \cdot \nabla) \rho = -\rho(\nabla \cdot \mathbf{v})$$

$$\rightarrow \frac{D\rho}{Dt} = -\rho(\nabla \cdot \mathbf{v})$$

Continuity equation



Momentum equation

- What accelerations act on a fluid element?

$$\mathbf{a} = -\frac{\nabla p}{\rho}$$

- This gives us the Euler equations

$$\frac{D\mathbf{v}}{Dt} = -\frac{\nabla p}{\rho}$$

or

$$\frac{\partial \rho \mathbf{v}}{\partial t} + \nabla \cdot (\rho \mathbf{v} \mathbf{v} + P \cdot \mathbf{I}) = 0$$



Momentum equation

- What accelerations act on a fluid element?

$$\mathbf{a} = -\frac{\nabla p}{\rho}$$

- This gives us the Euler equations

$$\frac{D\mathbf{v}}{Dt} = -\frac{\nabla p}{\rho}$$

or

$$\frac{\partial \rho \mathbf{v}}{\partial t} + \nabla \cdot (\rho \mathbf{v} \mathbf{v} + P \cdot \mathbf{I}) = 0$$

Your exercise!

Internal Energy equation

- From first law of thermodynamics,

$$d\epsilon = T \cancel{ds} - pdV, \quad dV = d\left(\frac{1}{\rho}\right) = \frac{-d\rho}{\rho^2}$$

$$d\epsilon \equiv d\left(\frac{u}{\rho}\right) = \frac{du}{\rho} - u \frac{d\rho}{\rho^2} = -pdV = p \frac{d\rho}{\rho^2}$$

Volume per unit mass

$$\frac{Du}{Dt} = \frac{u+p}{\rho} \frac{D\rho}{Dt} = \frac{\partial u}{\partial t} + (\mathbf{v} \cdot \nabla)u = -(u+p)\nabla \cdot \mathbf{v}$$

- Combine internal energy equation with the momentum equation

$$\frac{\partial \rho E}{\partial t} + \nabla \cdot [(\rho E + P)\mathbf{v}] = 0$$

Total Energy equation

$$E = \frac{1}{2}v^2 + \epsilon$$



Summary: Euler equations

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

Continuity equation

$$\frac{\partial \rho \mathbf{v}}{\partial t} + \nabla \cdot (\rho \mathbf{v} \mathbf{v} + P \cdot \mathbf{I}) = 0$$

Momentum equation

$$\frac{\partial \rho E}{\partial t} + \nabla \cdot [(\rho E + P) \mathbf{v}] = 0$$

Energy equation

ρ Gas density

\mathbf{v} Gas velocity

P Gas pressure

ϵ Gas specific internal energy

E Gas specific total energy

u Gas internal energy

$$u = \rho \epsilon$$

$$E = \frac{1}{2} v^2 + \epsilon$$



Summary: Euler equations + Gravity

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

Continuity equation

$$\frac{\partial \rho \mathbf{v}}{\partial t} + \nabla \cdot (\rho \mathbf{v} \mathbf{v} + P \cdot \mathbf{I}) = -\rho \nabla \Phi$$

Momentum equation

$$\frac{\partial \rho E}{\partial t} + \nabla \cdot [(\rho E + P) \mathbf{v}] = -\rho \mathbf{v} \cdot \nabla \Phi$$

Energy equation

$$\nabla^2 \Phi = 4\pi G \rho$$

Poisson equation

ρ Gas density

\mathbf{v} Gas velocity

P Gas pressure

ϵ Gas specific internal energy

E Gas specific total energy

u Gas internal energy

Φ Gravitational Potential

$$u = \rho \epsilon$$

$$E = \frac{1}{2} v^2 + \epsilon$$



Summary: Lagrangian form

$$\frac{D\rho}{Dt} = -\rho \nabla \cdot \mathbf{v}$$

Continuity equation

$$\frac{D\mathbf{v}}{Dt} = -\frac{\nabla p}{\rho} - \nabla \Phi$$

Momentum equation

$$\frac{Du}{Dt} = -(u + p) \nabla \cdot \mathbf{v}$$

Energy equation

ρ Gas density

\mathbf{v} Gas velocity

p Gas pressure

ϵ Gas specific internal energy

E Gas specific total energy

u Gas internal energy

Φ Gravitational Potential

$$u = \rho \epsilon$$

$$E = \frac{1}{2} v^2 + \epsilon$$



Equation of State (EoS)

Need an Equation of State to close system of equations.

For **ideal gas**,

$$P = (\gamma - 1)u = nk_{\text{B}}T$$

$$n = \rho/m \quad \text{Number density}$$

$$T \quad \text{Temperature}$$

$$k_{\text{B}} \quad \text{Boltzmann constant}$$

$$\gamma = C_v/C_p \quad \text{Adiabatic index}$$

- Isothermal EoS: $\gamma = 1$ $P = \rho c_s^2$
- Polytropic EoS: $P = K \rho^\gamma$

Finite Volume Method

		$U_{i,j+1}^n$		
	$U_{i-1,j}^n$	U_{ij}^n	$U_{i+1,j}^n$	
		$U_{i,j-1}^n$		

- Initial conditions of all hydrodynamics variables at time step $n=0$ (a given EoS)
- Boundary conditions at simulation boundaries
- Evolve hydrodynamics variables with a time step dt

$$\frac{\partial}{\partial t} \underbrace{\begin{pmatrix} \rho \\ \rho U \\ E \end{pmatrix}}_{W(x,t)} + \frac{\partial}{\partial x} \underbrace{\begin{pmatrix} \rho U \\ \rho U^2 + p \\ (E + p)U \end{pmatrix}}_{F(W)} = 0,$$

See Hsi-Yu Schive's lecture tomorrow

FLASH Code



FLASH Code: summary

- 3D AMR HD/MHD multi-physics code
- Mainly developed at FLASH center at University of Chicago (<http://flash.uchicago.edu>)
- Contribution from many individual groups
- Written principally in Fortran, with some C and Python (> 1 million lines; 25% are comments)
- Current release version 4.6.1
- > 700 scientists around the world
- > 1000 publications



FLASH Code: Requirements

- Fortran compiler
- MPI - Message Passing Interface
- HDF5 - Hierarchical Data Format



FLASH Code: research applications

- Thermonuclear flashes
- High energy density physics (HEDP)
- Fluid—structure interaction
- Star formation
- Star-star & Star-planetets interactions
- Core-collapse supernova
- Galaxy-galaxy cluster
- Magnetic field amplification
- Turbulence
- Cosmology
- ... (more)



FLASH Code: Physic solvers

- **Hydrodynamics**: unspilt PPM, WENO; split PPM; 2T + Radiation
- **Magneto-hydrodynamics (MHD)**: unspilt staggered mesh; split 8-wave
- **Equation of State (EoS)**: idea gas; degenerate ionized plasma; nuclear,..
- **Radiation Transfer**: multi-group flux-limited diffusion
- **Diffusion and Conduction**: implicit with AMR
- **Laser Energy Deposition**: geometric optics with inverse Bremsstrahlung
- **Opacity**: constant; multi-material tabular
- **Particles**: tracer; massive; sink; charged
- **Nuclear Burning**
- **Gravity**: constant; point mass; planar; self-gravity
- **Cosmology**
- **Heating and Cooling** as source terms
- **Magnetic Resistivity; Conductivity**
- **Primordial Chemistry**



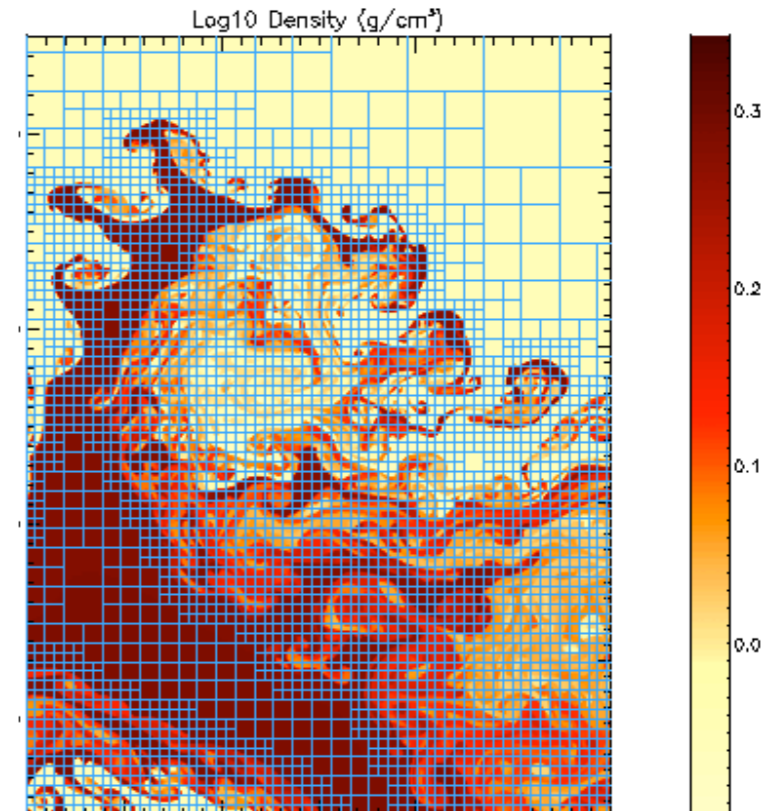
FLASH Code: Physic solvers

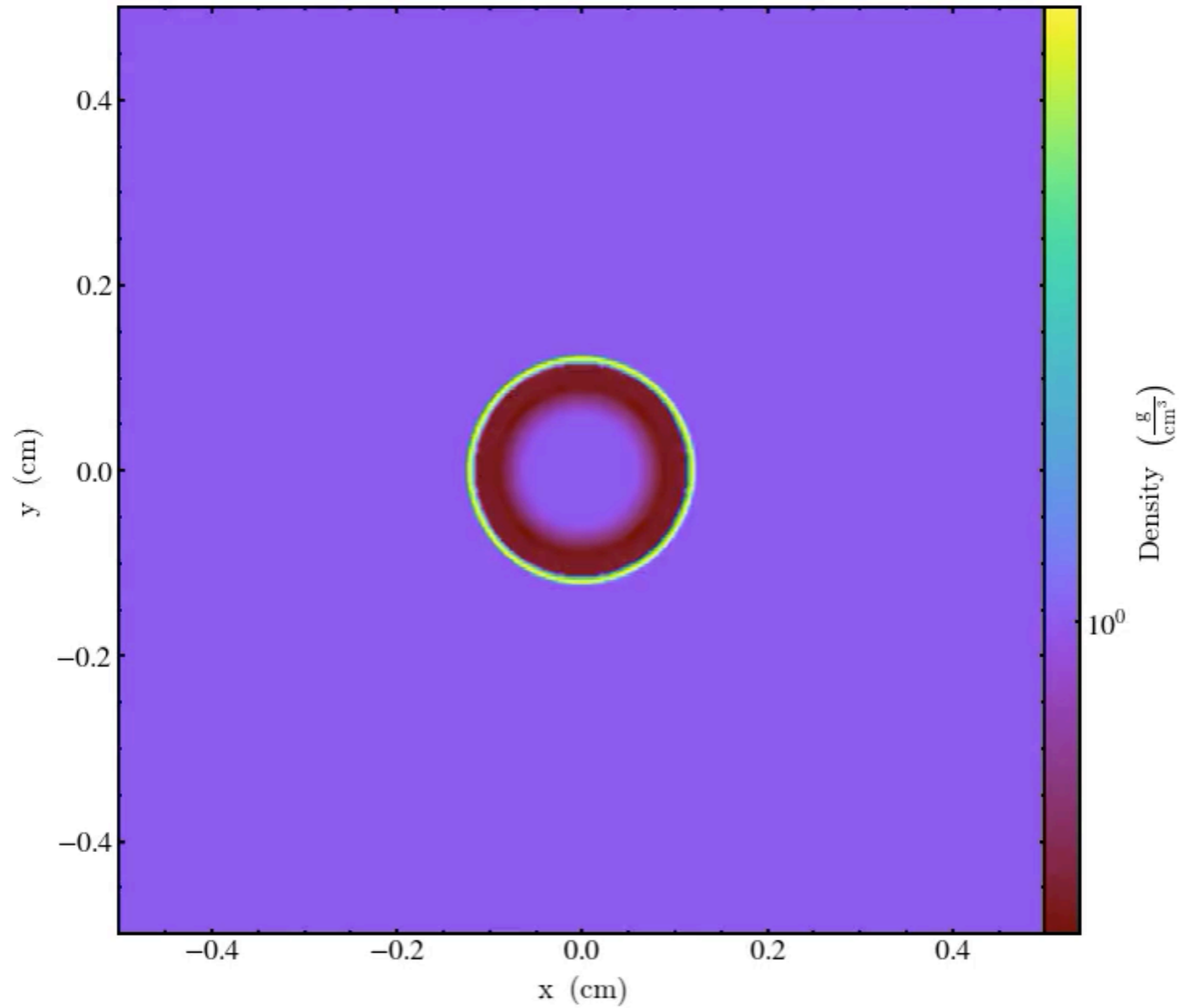
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- **Heating and Cooling** as source terms
- **Magnetic Resistivity; Conductivity**
- **Primordial Chemistry**

FLASH Code: Infrastructure

- **Driver**: split; unsplit
- **Grid**: uniform Grid; Adaptive Mesh Refinement (AMR)
- **GridParticles**: Lagrangian framework
- **GridSolvers**: multigrid; multipole; Barnes-Hut Tree; PFFT; direct solvers
- **IO**: Hierarchical Data Format 5 (HDF5); PnetCDF
- **Multispecies**
- **Runtime parameters**
- **Monitor**: MPI Timers; Hooks for TAU

Block structure PARAMESH AMR





See the afternoon lectures for data analysis and visualization



FLASH Code: research applications

- Highly modular (setup.py)

```

Driver/
Grid/
IO/
Multispecies/
Particles/
PhysicalConstants/
RuntimeParameters/
Simulation/
diagnostics/
flashUtilities/
monitors/
multiprocessorTools/
numericalTools/
physics/

```

```

Cosmology/
Diffuse/
Eos/
Gravity/
Hydro/
ImBound/
IncompNS/
RadTrans/
RayTrace/
SolidMechanics/
TreeRay/
materialProperties/
sourceTerms/
utilities/

```

```

Burn/
Cool/
Deleptonize/
EnergyDeposition/
Flame/
Heat/
Heatexchange/
Ionize/
Polytrope/
PrimordialChemistry/
Stir/
Turb/

```


FLASH Code: Simulation Domain

- AMR

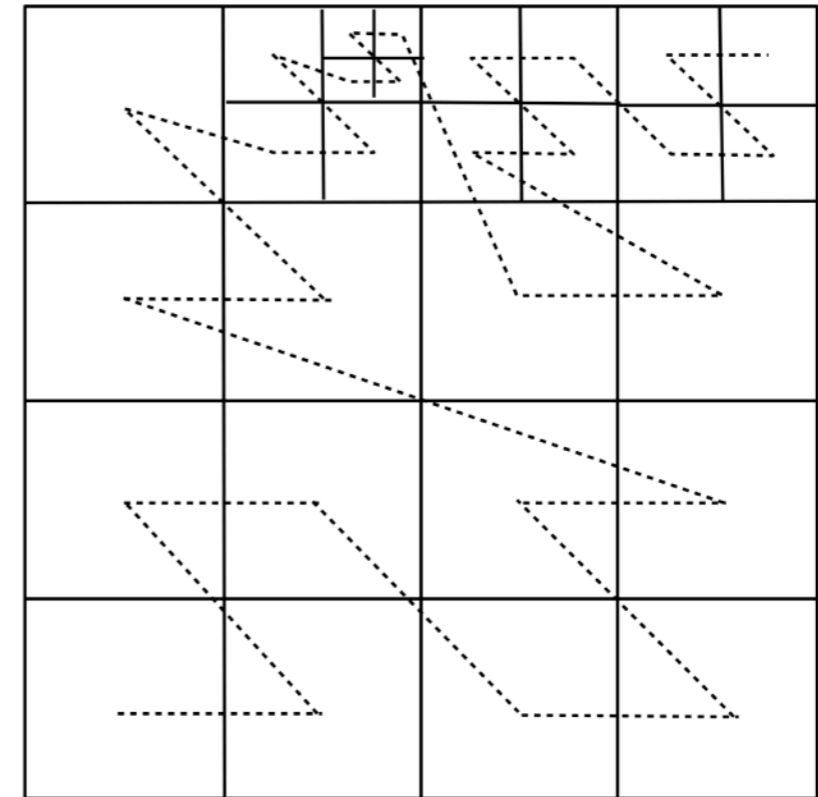
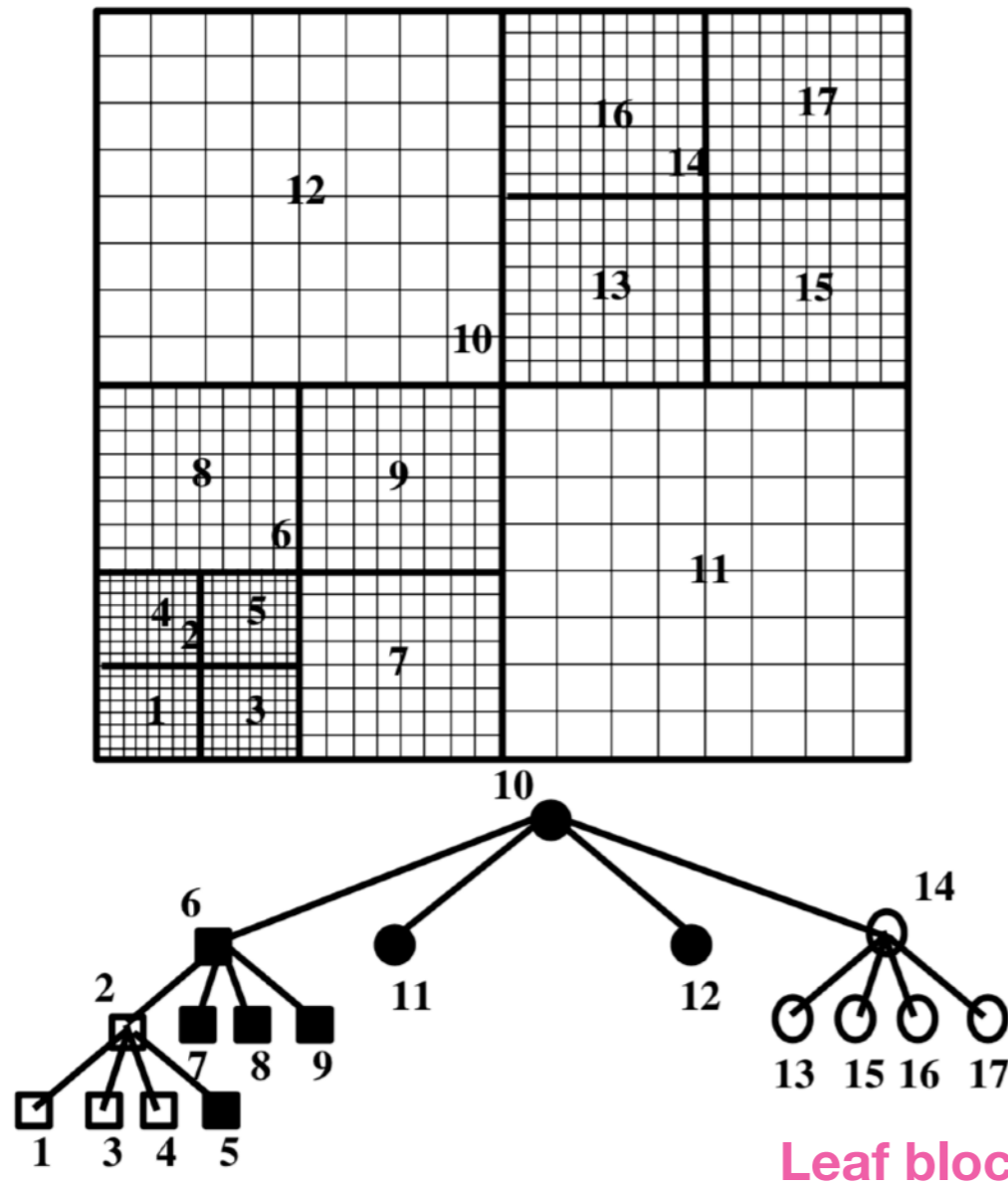
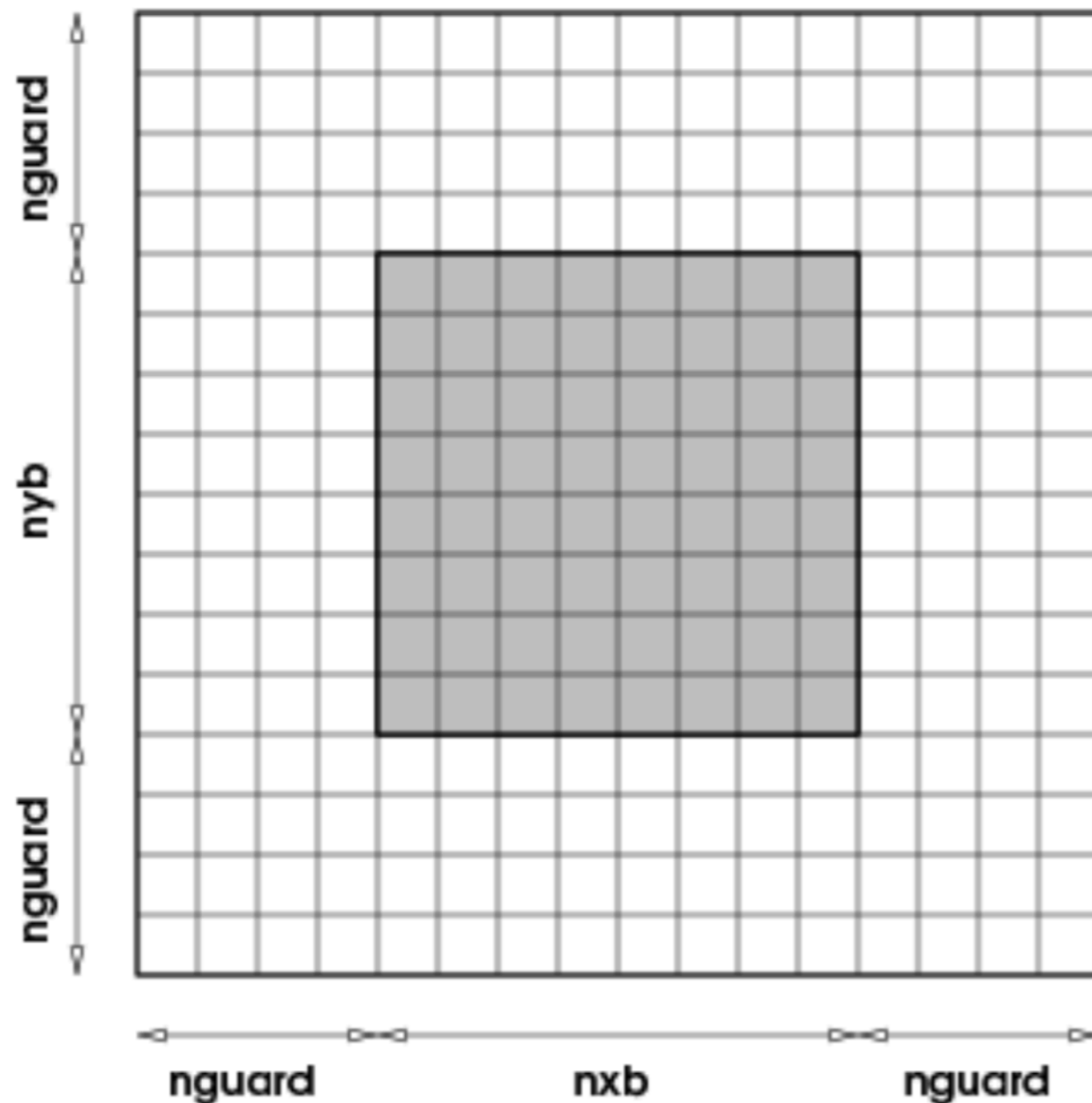


FIG. 3.—Morton space-filling curve for an arbitrary set of blocks of differing spatial resolution

FLASH Code: Simulation Domain

- Blocks



- The grid is composed of blocks
- FLASH4: all blocks are of same size
- May cover different fractions of the physical domain, depending on a block's resolution
- Data storage area for each block reserves space for some layers of guard cells
- Size of guard cells can be changed by putting a line "GUARDCELLS 6" in the Config file. Default is 4.



FLASH Code: Variables

- Density: DENS_VAR
- Velocities: VELX_VAR, VELY_VAR, VELZ_VAR
- Temperature: TEMP_VAR
- Pressure: PRES_VAR
- Specific internal energy: EINT_VAR



Simulations

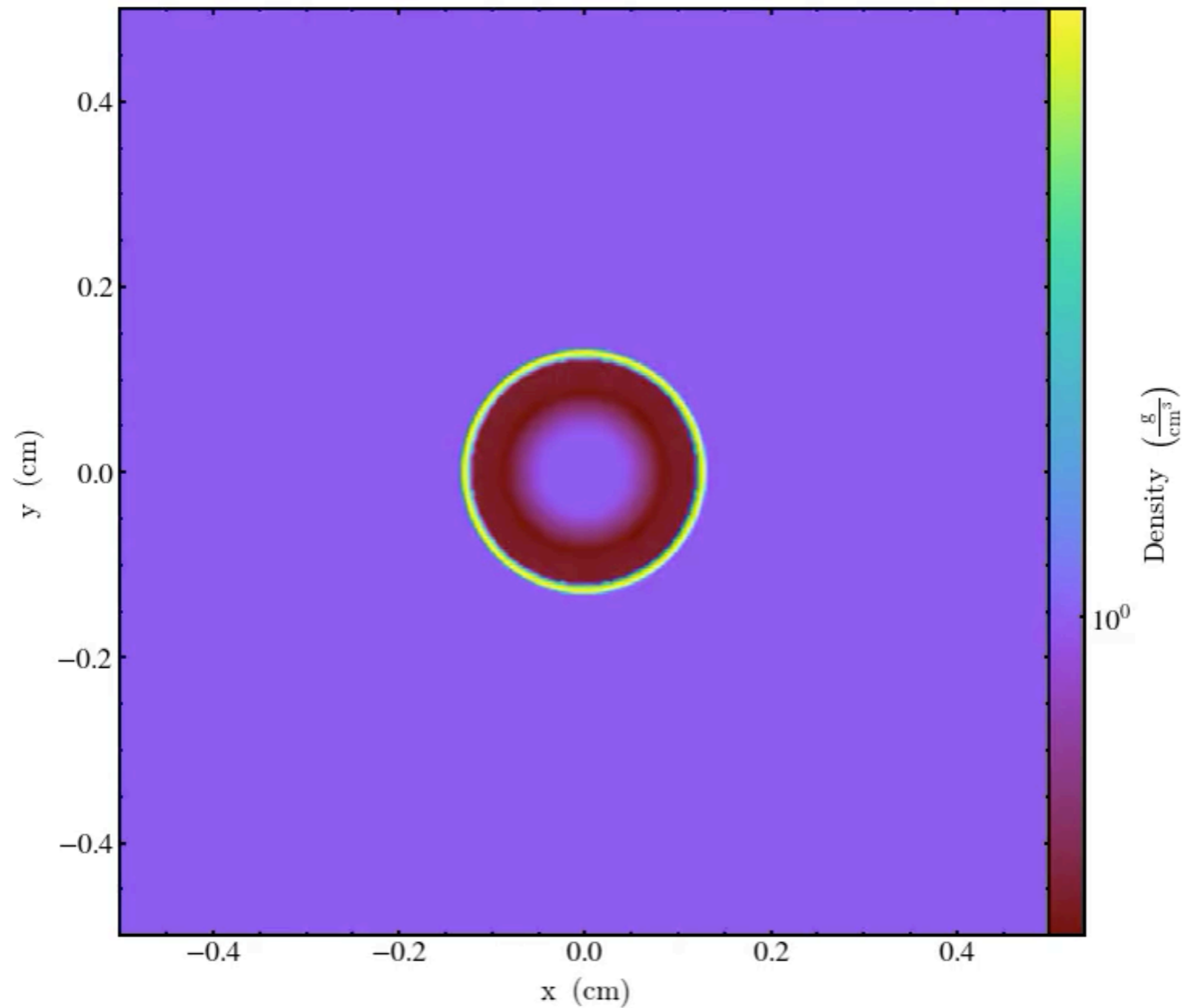
- All sample simulations can be found in
FLASH/source/Simulation/SimulationMain/
- A simulation requires the following 5 files:
 - Config
 - Makefile:
 - Simulation_data.F90
 - Simulation_init.F90
 - Simulation_initBlock.F90



Design a FLASH simulation

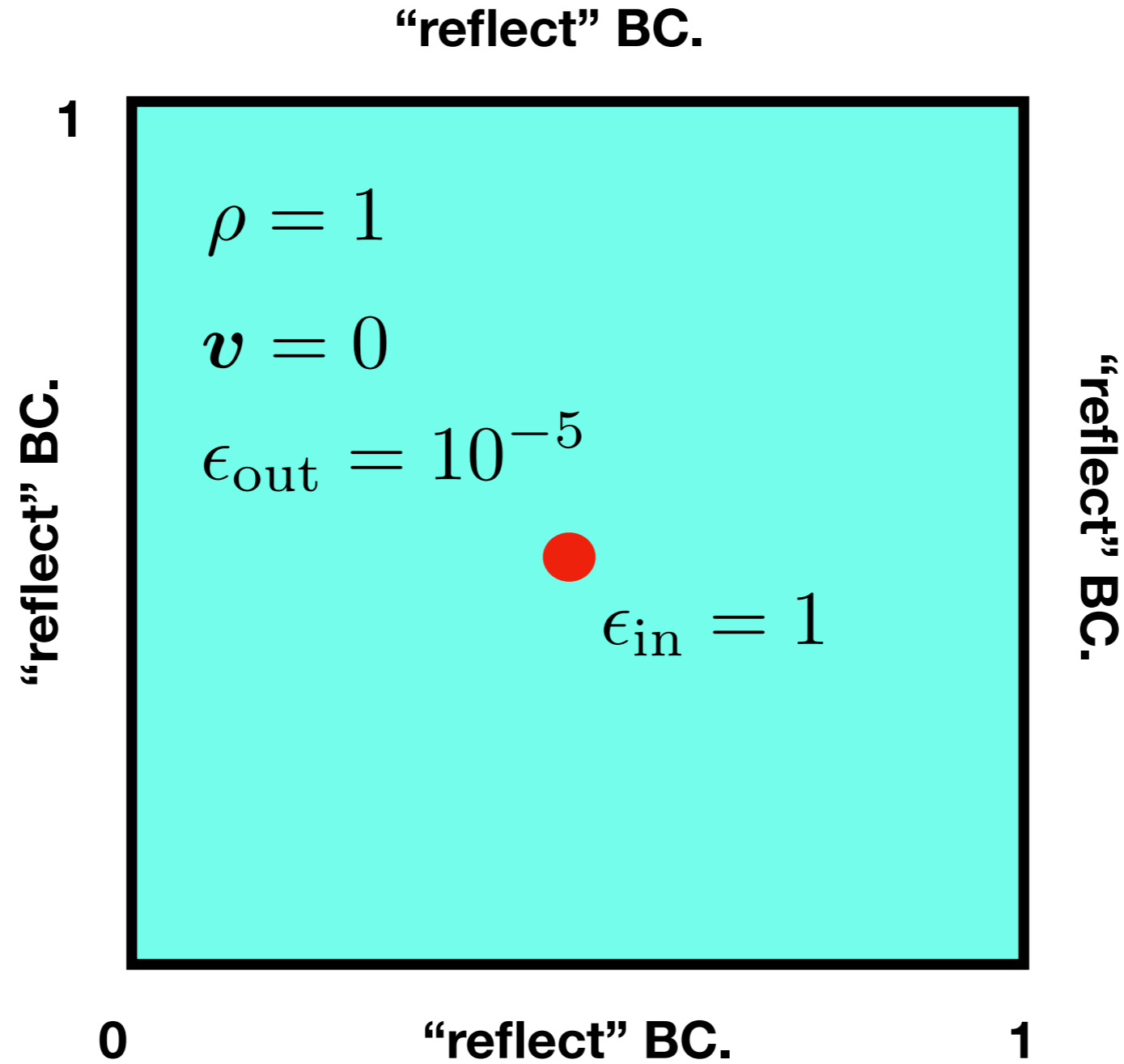
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 - Config
 - Makefile:
 - Simulation_data.F90
 - Simulation_init.F90
 - Simulation_initBlock.F90

A simple explosion simulation





The initial condition



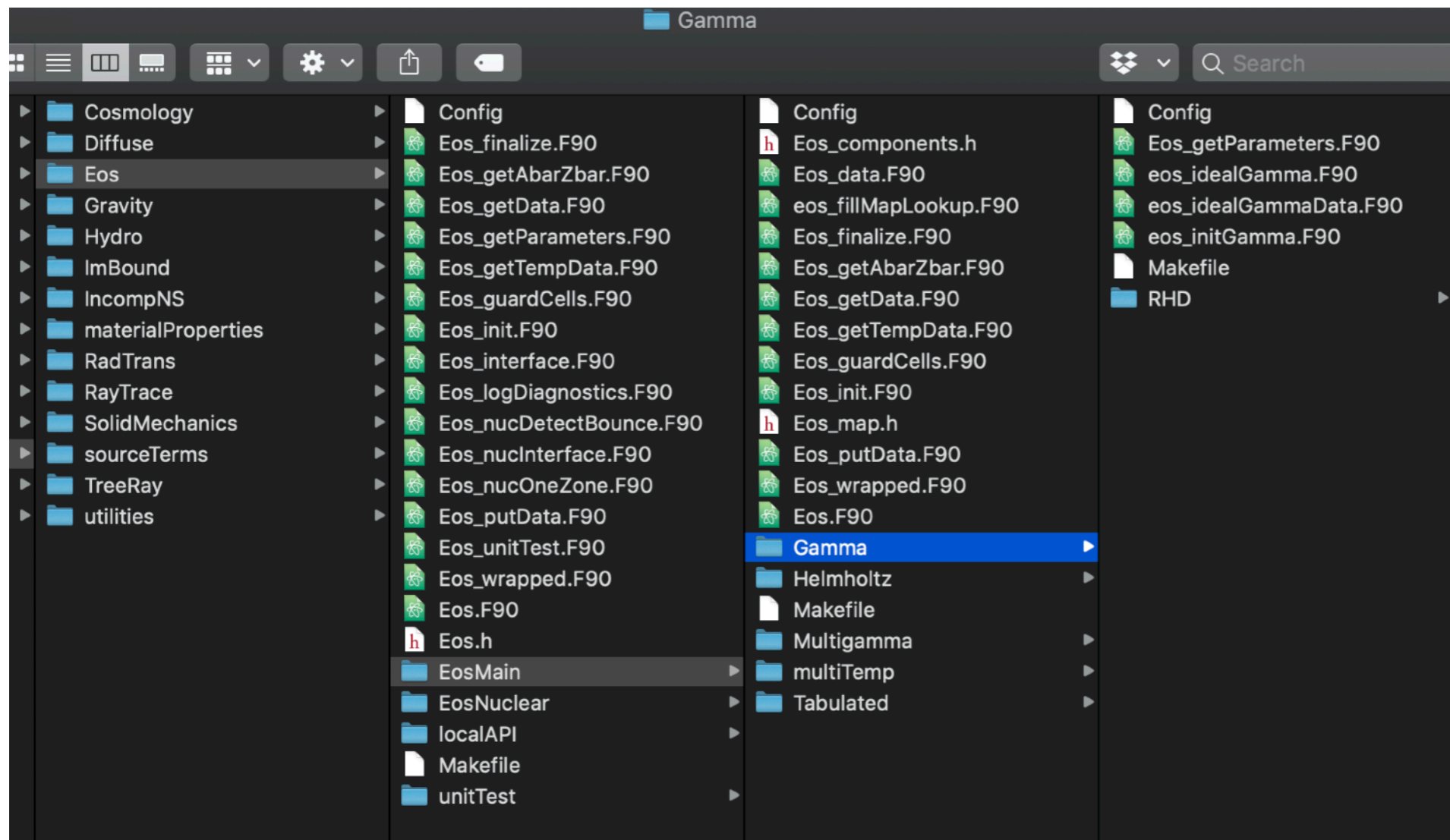


Config

```
Config + (~/codes/FLASH/s...ationMain/Template) - VIM
1 # The Configuration file for a test problem
2
3 REQUIRES Driver
4 REQUIRES physics/Hydro
5 REQUIRES physics/Eos
6
7 D sim_rho0 The ambient density [g/cc]
8 PARAMETER sim_rho0 REAL 1. [0 to ]
9
10
11 █
~
~
~
Config [+] 11,0-1 All
pan vega.astr.nthu.edu.tw ~/c/F/s/S/S/Template vim - zsh 8/28, 3:16 PM
```


Using a Flash Unit

- REQUIRES physics/EoS
- Default: physics/Eos/EosMain/Gamma





Simulation_data.F90

```
25 module Simulation_data
26 #include "Flash.h"
27   implicit none
28
29   !! *** Runtime Parameters *** !!
30   real, save :: sim_rho0
31
32   !! *** Variables pertaining to Simulation
33   logical, save :: sim_gCell
34
35   integer, save :: sim_meshMe
36 end module Simulation_data
37
```



Simulation_init.F90

```
28 subroutine Simulation_init()
29
30 use Simulation_data
31 use Driver_interface, ONLY : Driver_getMype, Driver_abortFlash
32 use RuntimeParameters_interface, ONLY : RuntimeParameters_get
33 use Logfile_interface, ONLY : Logfile_stamp
34 implicit none
35 #include "constants.h"
36 #include "Flash.h"
37 #include "Multispecies.h"
38
39 call Driver_getMype(MESH_COMM, sim_meshMe)
40
41 call RuntimeParameters_get('sim_rho0', sim_rho0)
42
43 call Logfile_stamp( "initializing the Template problem", &
44                   "[Simulation_init]")
45
46 end subroutine Simulation_init
```



Simulation_init.F90

```
28 subroutine Simulation_init()
29
30 use Simulation_data
31 use Driver_interface, ONLY : Driver_getMype, Driver_abortFlash
32 use RuntimeParameters_interface, ONLY : RuntimeParameters_get
33 use Logfile_interface, ONLY : Logfile_stamp
34 implicit none
35 #include "constants.h"
36 #include "Flash.h"
37 #include "Multispecies.h"
38
39 call Driver_getMype(MESH_COMM, sim_meshMe)
40
41 call RuntimeParameters_get('sim_rho0', sim_rho0)
42
43 call Logfile_stamp( "initializing the Template problem", &
44                   "[Simulation_init]")
45
46 end subroutine Simulation_init
```



Simulation_initBlock.F90

```

30 !!***
31
32 subroutine Simulation_initBlock(blockID)
33
34 #include "constants.h"
35 #include "Flash.h"
36 #include "Eos.h"
37
38 use Simulation_data, ONLY : sim_rho0
39
40 use Grid_interface, ONLY : Grid_getBlkIndexLimits, &
41   Grid_getCellCoords, Grid_getBlkPtr, Grid_releaseBlkPtr
42 use Eos_interface, ONLY : Eos, Eos_wrapped
43
44
45 implicit none
46
47 integer, intent(in) :: blockID
48
49
50 !! Handling AMR blocks and pointer
51
52 integer :: i, j, k, n
53 real, dimension(:,:,:), pointer :: solnData
54 integer, dimension(2,MDIM) :: blkLimits, blkLimitsGC
55 integer :: sizeX, sizeY, sizeZ
56 logical :: gcell = .true.
57
58 !! Physical coordinates
59 real :: xx, yy, zz
60 real :: radius
61 real, allocatable, dimension(:) :: yCenter, xCoord, zCoord
62
63 !! physical parameters
64 real :: rho, pres, velx, vely
65

```

```

66 ! get the integer index information for the current block
67 call Grid_getBlkIndexLimits(blockId, blkLimits, blkLimitsGC)
68 call Grid_getBlkPtr(blockId, solnData)
69
70
71 sizeX = blkLimitsGC(HIGH, IAXIS)
72 sizeY = blkLimitsGC(HIGH, JAXIS)
73 sizeZ = blkLimitsGC(HIGH, KAXIS)
74 allocate(yCenter(sizeY))
75 allocate(xCoord(sizeX))
76 allocate(zCoord(sizeZ))
77 yCenter = 0.0
78 xCoord = 0.0
79 zCoord = 0.0
80
81 if (NDIM == 3) call Grid_getCellCoords&
82   (KAXIS, blockId, CENTER, gcell, zCoord, sizeZ)
83 if (NDIM >= 2) call Grid_getCellCoords&
84   (IAXIS, blockId, CENTER, gcell, xCoord, sizeX)
85
86 call Grid_getCellCoords(JAXIS, blockId, CENTER, gcell, yCenter, sizeY)
87

```



Simulation_initBlock.F90

```
95 do k = blkLimits(LOW,KAXIS),blkLimits(HIGH,KAXIS)
96
97 ! get the coordinates of the cell center in the z-direction
98 zz = zCoord(k)
99 do j = blkLimits(LOW,JAXIS),blkLimits(HIGH,JAXIS)
100 ! get the coordinates of the cell center in the y-direction
101 yy = yCenter(j)
102 ! The position in the current yz-row.
103 do i = blkLimits(LOW,IAXIS),blkLimits(HIGH,IAXIS)
104
105 ! get the cell center, left, and right positions in x
106 xx = xCoord(i)
107
108 radius = sqrt(xx**2 + yy**2 + zz**2)
109
110 if (radius .le. 0.1) then
111 pres = (4./3. - 1.)*1.0
112 else
113 pres = 1.e-5
114 endif
115 rho = sim_rho0
116 velx = 0.0
117 vely = 0.0
118
119 solnData(DENS_VAR, i,j,k) = rho
120 solnData(PRES_VAR, i,j,k) = pres
121 solnData(VELX_VAR, i,j,k) = velx
122 solnData(VELY_VAR, i,j,k) = vely
123 solnData(VELZ_VAR, i,j,k) = 0.0
124
125 enddo
126 enddo
127 enddo
128
129 call Grid_releaseBlkPtr(blockID, solnData)
130 call Eos_wrapped(MODE_DENS_PRES,blkLimits,blockId)
131
132 deallocate(yCenter)
133 deallocate(xCoord)
134 deallocate(zCoord)
135
136 return
137 end subroutine Simulation_initBlock
```



Setup the Simulation

'FLASH/source/Simulation/SimulationMain/Template

- A machine depend Makefile.h

```
~/codes/FLASH/sites(sim_school*) » ls
Aliases                buckbeak                ellipse09.uchicago.edu  hyades.ucsc.edu        mhd2.asci.uchicago.edu  scooter.asci.uchicago.edu
FomalhautIMP           cetus.asci.uchicago.edu  ellipse10.uchicago.edu  hydra.si.edu           midway2.rcc.uchicago.edu  seaborg.nersc.gov
FomalhautMBP4         clark.asci.uchicago.edu  ellipse11.uchicago.edu  icc-9.0_fornax.uchicago.edu  mira.alcf.anl.gov         skeeter.asci.uchicago.edu
P640278.nist.gov      clogin1                 ellipse_pgf.uchicago.edu  ignition                mongchi.soe.ucsc.edu      sphere.uchicago.edu
Prototypes            code.uchicago.edu       eugenia.asci.uchicago.edu  intel-hedp             mongchi.uchicago.edu     splash.seas.gwu.edu
SEAS10926.gwu.edu     coyote.lanl.gov          eureka.alcf.anl.gov      intel-mpi2             myristica.asci.uchicago.edu  splash.seas.gwu.edu_KPD
SEAS10927.gwu.edu     crash2.umd.edu           fen.bluegene.bnl.gov     intel-ompi             nag-mpi1                  sunspot.uchicago.edu
SEAS10927UB.gwu.edu  cthinkpad               fenp.bluegene.bnl.gov     intrepid.alcf.anl.gov  nag-mpi2                  surveyor.alcf.anl.gov
SEAS12982.gwu.edu     ctsv.astro.sunysb.edu   flash.uchicago.edu      jacquard.nersc.gov     nagini.uchicago.edu      tp-login2
SEAS13033.gwu.edu     cube.uchicago.edu       flashviz.uchicago.edu   jaguar.ccs.ornl.gov    oakley.osc.edu            tsoodzil.astro.uiuc.edu
absoft-mpi2           datastar.sdsc.edu        fleetwood.astro.sunysb.edu  jaguar.nccs.gov        optix.cs.uoregon.edu       tuxedo.uchicago.edu
alc.llnl.gov          depththought.umd.edu_tfitz  fomalhaut                 jubl.zam.kfa-juelich.de  osel.uchicago.edu        uffda.asci.uchicago.edu
amazon.nswccd         duce.gsfc.nasa.gov       fornax.uchicago.edu     karloff.lbl.gov        p655-4.nic.uoregon.edu     variable.as.arizona.edu
animal5              edison.nersc.gov         franklin.nersc.gov        khorba.uchicago.edu  pg-mpi2-32                variable.ph.ua.edu
archimedes.uchicago.edu  elan.uchicago.edu      fusion.lcrc.anl.gov      klaus-laptop           purple.llnl.gov            vega.astr.nthu.edu.tw
asterix.asci.uchicago.edu  eldorado.astro.sunysb.edu  fusion.lcrc.uchicago.edu  kraken.nics.tennessee.edu  pyramid.uchicago.edu     vestalac1
babbage.nersc.gov      eldorado.uchicago.edu   gin.asci.uchicago.edu   lahey-mpi2             qsc.lanl.gov              watanlsn.watson.ibm.com
bassi.nersc.gov        ellipse.uchicago.edu     gnu-ins                  lenovolaptop           r1.oit.ua.edu             zeus.llnl.gov
beagle.ci.uchicago.edu  ellipse02.uchicago.edu  gnu-mpi2                 liturchi.uchicago.edu  ramsusii.mps.ohio-state.edu  zingiber.asci.uchicago.edu
bgl.llnl.gov           ellipse03.uchicago.edu  gnu-ompi                 login1.pads.ci.uchicago.edu  ranger.tacc.utexas.edu     zingiber.uchicago.edu
bgl.mcs.anl.gov        ellipse04.uchicago.edu  hawkmoon.uchicago.edu   login2.pads.ci.uchicago.edu  rc2.ua.edu                redstorm.sandia.gov
bgl.sdsc.edu           ellipse05.uchicago.edu  hera.llnl.gov            lupin.uchicago.edu     saguaro.fulton.asu.edu     scarf.rl.ac.uk
bonsai.cfa.harvard.edu  ellipse06.uchicago.edu  hopper.nersc.gov         macbro.uchicago.edu
brassica.asci.uchicago.edu  ellipse08.uchicago.edu  hpc.msu.edu              maitreyi
```




Setup the Simulation

- Run a setup script

A simple example

```
./setup Template -2d -auto
```

```
./setup Template -2d -auto -maxblocks=2000 -nxb=8 -nyb=8|
```

A more complicated example

```
~/codes/FLASH(sim_school*) » ./setup CoreCollapse/IDSA -auto -2d +cylindrical -objdir ccsn2dIdsa -nxb=16 -nyb=16 -max  
blocks=1400 threadBlockList=False +pm4dev threadWithinBlock=False +newMpole --with-unit=source/physics/RadTrans/RadTr  
ansMain/IDSA/rk2█
```



Setup the Simulation

```
~/codes/FLASH(sim_school*) » ./setup  
Processing Shortcut file: /Users/pan/codes/FLASH/bin/setup_shortcuts.txt  
usage: setup <problem-name> [options] [VAR=VALUE]...
```

problem-name: see source/Simulation/SimulationMain directory
options:

(Science Options)

```
-auto -[123]d  
-maxblocks=<#> -nxb=<#> -nyb=<#> -nzb=<#>  
-with-unit=<unit> -with-library=<libname>[,args]  
-without-unit=<unit> -without-library=<libname>
```

(Setup and Make Options)

```
-verbose=[DEBUG|INFO|WARN|IMPINFO|ERROR]  
[-site=<site> | -ostype=<ostype>]  
-makefile=<extension>  
[-optl -debug | -test ]  
-objdir=<relative obj directory>  
-defines=<defines> -unitsfile=<filename>  
-datafiles=<wildcard> -parfile=<filename>  
-fbs -nofbs -tau=<makefile>
```

(Misc Options)

```
-makehide -noclobber -portable -help
```

- * For GNU compatibility, options may be prefixed by -- instead of - as well
- * -unit and -library are considered equivalent to -with-unit and -with-library respectively.
- * For information regarding the [VAR=VALUE] options and using 'setup variables' refer to User's Guide.
- * To read how shortcuts work see README.shortcuts in your bin directory



Setup shortcuts

To use a shortcut add '+shortcut' to your setup line.
For example ./setup Sod -auto +ug

3t	ThreeT=1 --defines=FLASH_3T
3tr	+3t RadFlahThreeT=True
8wave	--with-unit=physics/Hydro/HydroMain/split/MHD_8Wave +grid --gridinterpolation=native
asynclaser	+laser useAsynLaser=True --without-unit=physics/sourceTerms/EnergyDeposition/EnergyDeposition
Main/Laser/LaserIO	
cartesian	--geometry=cartesian
chombo_amr	--unit=Grid/GridMain/Chombo/AMR --index-reorder Grid=Chombo --nofbs --makefile=chombo chomboCo
compatibleHydro=True	
chombo_ug	--unit=Grid/GridMain/Chombo/UG --index-reorder Grid=Chombo --maxblocks=1 --nofbs --makefile=ch
chombo chomboCompatibleHydro=True	
cube16	--nxb=16 --nyb=16 --nzb=16
cube32	--nxb=32 --nyb=32 --nzb=32
cube64	--nxb=64 --nyb=64 --nzb=64
curv-pm2	+pm2 --unit=Grid/GridMain/paramesh/Paramesh2 --with-unit=Grid/GridMain/paramesh/Paramesh2/mono
tonic	
curvilinear	--curvilinear
cylindrical	--geometry=cylindrical
ddt	use_ddt=True --defines=DDT --with-unit=source/monitors/Debugger/DebuggerMain/dmalloc
default	--with-library=mpi +io +grid --gridinterpolation=monotonic FlashAvoidOrrery=True
dmalloc	use_dmalloc=True --defines=DMALLOC --with-unit=source/monitors/Debugger/DebuggerMain/dmalloc
dynamicmem-pm3	+dynamicmem-pm40
dynamicmem-pm40	+pm40 ParameshLibraryMode=True
gravmgrid	+pm4dev --with-unit=physics/Gravity/GravityMain/Poisson/Multigrid
gravmpole	--with-unit=physics/Gravity/GravityMain/Poisson/Multipole
gravpfftnofbs	+ug +nofbs --with-unit=physics/Gravity/GravityMain/Poisson/Pfft
grid	--unit=Grid
hdf5	I0=hdf5
hdf5typeio	+io +parallelI0 +hdf5 typeI0=True
io	--with-unit=I0
laser	--unit=physics/sourceTerms/EnergyDeposition/EnergyDepositionMain/Laser --without-unit=Particle



Setup shortcuts (continues)

```

newmpole          +noMgrid +noDefaultMpole +gravMpole --with-unit=Grid/GridSolvers/Multipole_new
nodefaultmpole   --without-unit=Grid/GridSolvers/Multipole
nofbs            --nofbs +ug parallelIO=True
noio              --without-unit=physics/sourceTerms/EnergyDeposition/EnergyDepositionMain/Laser/LaserIO --witho
ut-unit=IO
nolog            --without-unit=monitors/Logfile
nomgrid          --without-unit=physics/Gravity/GravityMain/Poisson/Multigrid
npg              npg=True
parallelio       parallelIO=True
pic              +ug --unit=Grid/GridParticles/GridParticlesMove --without-unit=Grid/GridParticles/GridParticle
sMove/UG --without-unit=Grid/GridParticles/GridParticlesMove/UG/Directional
pipeline         --unit=multiprocessorTools/Pipeline
pm2              +grid Grid=PM2
pm3              +pm40
pm40             +grid Grid=PM40
pm4dev           +pm4dev_basic FlashAvoidOrrery=True
pm4dev_basic    +grid Grid=PM4DEV
pm4dev_clean    +pm4dev_basic FlashAvoidOrrery=False
pm4dev_fixed    +pm4dev ParameshLibraryMode=False
pnetcdf         IO=pnetcdf
pnettypeio      +io +parallelIO +pnetcdf typeIO=True
polar           --geometry=polar
protonemission  --unit=diagnostics/ProtonEmission --without-unit=Particles
protonimaging   --unit=diagnostics/ProtonImaging --without-unit=Particles
ptdens         --without-unit=Particles/ParticlesInitialization/Lattice --without-unit=Particles/ParticlesIni
tialization/WithDensity/CellMassBins --unit=Particles/ParticlesMain --unit=Particles/ParticlesInitialization/WithDensity --pa
rticlemethods=TYPE=passive,INIT=With_Density
ptio            +ug --with-unit=Particles
purehydro       physicsMode=hydro
rnf             --3d --nxb=8 --nyb=16 --nzb=32 --nofbs +ug
serialio        parallelIO=False
spherical       --geometry=spherical
spherical-pm2   +pm2 +spherical
splithydro      --with-unit=physics/Hydro/HydroMain/split --without-unit=physics/Hydro/HydroMain/unsplit Split
Driver=True
supportppmupwind SupportPpmUpwind=True

```



Compile source codes

- An object/ folder will be generated

```
cd object
```

```
make
```

```
mpi_unpack_edges.o mpi_unpack_fluxes.o mpi_unpack_tree_info.o mpi_wrapper_double.o mpi_wrapper_int
mpi_wrapper_real.o nameSyntaxError.o nameValueLL_bcast.o nameValueLL_data.o nameValueLL_get.o na
alueLL_getNum.o nameValueLL_logRules.o nameValueLL_rules.o nameValueLL_set.o paramesh_comm_data.o
aramesh_interfaces.o paramesh_mpi_interfaces.o pc_interface.o pc_utilities.o physicaldata.o pois
ist.o prolong_arrays.o rationalize_fetch_list.o removeNullChar.o rp_getArgument.o rp_getOpt.o rp
s.o rp_storeIgnoredParams.o send_block_data.o setup_buildstamp.o setup_buildstats.o setup_flashRe
its.o timings.o tmr_buildSummary.o tmr_create.o tmr_etime.o tmr_findTimerIndex.o tmr_getMaxCallSt
erParents.o tmr_init.o tmr_lookupIndex.o tmr_stackLib.o tree.o tree_search_for_surrblks.o umap.o
constRealZeroFn1.o ut_conversionInterface.o ut_convertToArrayIndicies.o ut_convertToMemoryOffset
Interface.o ut_getFreeFileUnit.o ut_hunt.o ut_interpolationInterface.o ut_parabolicInterpol.o ut
.o ut_qsort.o ut_qsortC.o ut_qsortInterface.o ut_quadraticInterpol.o ut_sortInterface.o ut_sortOnf
t_sysMemInterface.o ut_sysMemStats.o ut_sysMemSummaryStats.o ut_sys_mem_usage.o workspace.o -L/Us
1-gnu/lib -lhdf5 -lhdf5_fortran -lz -DH5_USE_16_API
SUCCESS
```



Run Flash

- Requires

the executable file `flash4` and

the runtime parameter file `flash.par`



Runtime Parameters (flash.par)

```
1 # Runtime parameters for the Kelvin-Helmh
2
3 # Parameters for initial model
4
5 sim_rho0 = 1.0
6
7
8 # Grid dimensionality and geometry
9
10 geometry = cartesian
11
12 # Size of computational volume
13
14 xmin      = -0.5
15 xmax      = 0.5
16 ymin      = -0.5
17 ymax      = 0.5
18
```



Runtime Parameters

```
19 #           Boundary conditions
20
21 xl_boundary_type = "reflect"
22 xr_boundary_type = "reflect"
23
24 yl_boundary_type = "reflect"
25 yr_boundary_type = "reflect"
26
27
28 #           Simulation (grid, time, I/O) parameters
29
30 cfl           = 0.8
31 basenm       = "my_sim_"
32 restart      = .false.
33
```




Runtime Parameters

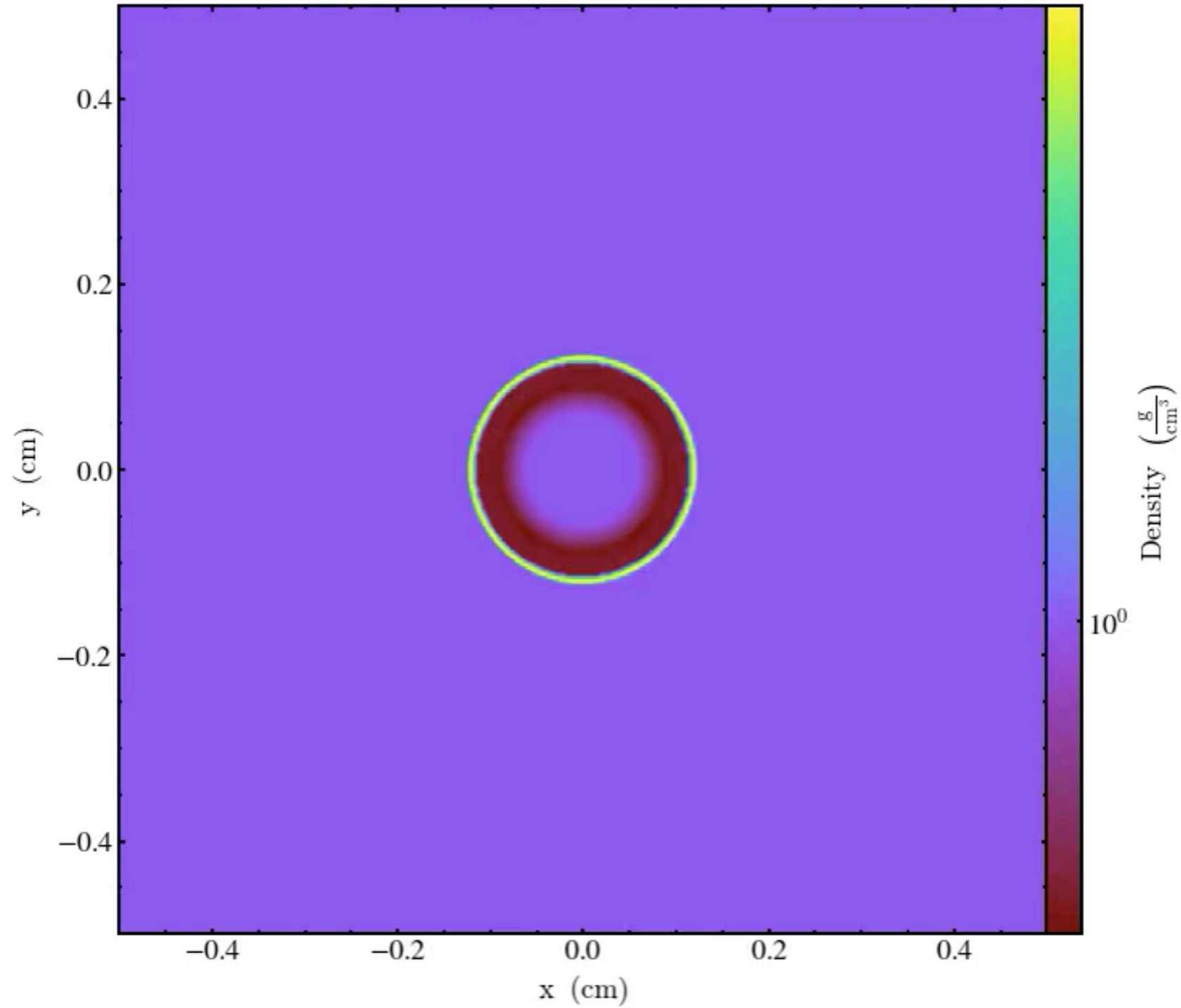
```
34 # checkpoint file output parameters
35 checkpointFileIntervalTime = 0.5
36 checkpointFileIntervalStep = 0
37 checkpointFileNumber      = 0
38
39 # plotfile output parameters
40 plotfileIntervalTime = 0.1
41 plotfileIntervalStep = 0
42 plotfileNumber       = 0
43
44 nend                = 1000000
45 tmax                = 10.0
46
47 run_comment        = "my first flash simulation"
48 log_file           = "my_sim.log"
49 eintSwitch         = 1.e-4
50
51
52 plot_var_1 = "dens"
53 plot_var_2 = "pres"
54 plot_var_3 = "temp"
55 plot_var_4 = "velx"
56 plot_var_5 = "vely"
--
```

```
59 # AMR refinement parameters
60 lrefine_min      = 1
61 lrefine_max      = 6
62 refine_var_1    = "dens"
63 refine_var_2    = "pres"
64
```



Outputs

- my_sim.log
- my_sim.dat
- my_sim_hdf5_chk_*
- my_sim_hdf5_plt_cnt_*



See the afternoon lectures for data analysis and visualization



Runtime parameters documentation

- [http://flash.uchicago.edu/site/flashcode/
user_support/rpDoc_4p61.py](http://flash.uchicago.edu/site/flashcode/user_support/rpDoc_4p61.py)

RUNTIME PARAMETERS DOCUMENTATION FOR FLASH RELEASE 4.6.1

All Chapters

Burn
Conductivity
Cool
Cosmology
Debugger
Deleptonize
Diffuse
Driver
EnergyDeposition
Eos
Flame
Gravity
Grid
Heat
Heatexchange
Hydro
IO
IncompNS
Ionize
Logfile
MagneticResistivity
MassDiffusivity
Opacity
Particles
PhysicalConstants
PlasmaState
Polytrope
PrimordialChemistry
Profiler
ProtonEmission
ProtonImaging
RadTrans
RayTrace
RungeKutta
Simulation
Stir
ThomsonScattering
Timers
TreeRay
Turb
Viscosity
XrayImaging



FLASH API

- http://flash.uchicago.edu/site/flashcode/user_support/robodoc-FLASH4_4p61/

FLASH4.6.1 API

ROBODoc 4.99 manual

Generated from /asc/asci2/site/flashcode/secure/release_4p6/ with ROBODoc v4.99.8 on Mon Sep 02 01:17:10 2019

[Sourcefiles] [Functions] [Modules]

FUNCTIONS

diagnostics/ProtonEmission/ProtonEmission	Particles/Particles_finalize
diagnostics/ProtonEmission/ProtonEmission_finalize	Particles/Particles_getCountPerBlk
diagnostics/ProtonEmission/ProtonEmission_init	Particles/Particles_getGlobalNum
diagnostics/ProtonImaging/ProtonImaging	Particles/Particles_getLocalNum
diagnostics/ProtonImaging/ProtonImaging_finalize	Particles/Particles_init
diagnostics/ProtonImaging/ProtonImaging_init	Particles/Particles_initData
diagnostics/ThomsonScattering/ThomsonScattering	Particles/Particles_initForces
diagnostics/ThomsonScattering/ThomsonScattering_finalize	Particles/Particles_initPositions
diagnostics/ThomsonScattering/ThomsonScattering_init	Particles/Particles_longRangeForce
diagnostics/XrayImaging/XrayImaging	Particles/Particles_manageLost
diagnostics/XrayImaging/XrayImaging_finalize	Particles/Particles_mapFromMesh
diagnostics/XrayImaging/XrayImaging_init	Particles/Particles_putLocalNum
Driver/Driver_abortFlash	Particles/Particles_sendOutputData
Driver/Driver_abortFlashC	Particles/Particles_shortRangeForce
Driver/Driver_checkMPIErrorCode	Particles/Particles_sinkAccelGasOnSinksAndSinksOnGas
Driver/Driver_computeCellLocations	Particles/Particles_sinkAdvanceParticles
Driver/Driver_computeDt	Particles/Particles_sinkComputeDt
Driver/Driver_diagnostics	Particles/Particles_sinkCreateAccrete
Driver/Driver_driftBlock	Particles/Particles_sinkInit
Driver/Driver_driftSetSrcLoc	Particles/Particles_sinkMarkRefineDerefine
Driver/Driver_driftUnk	Particles/Particles_sinkMoveParticles
Driver/Driver_evolveFlash	Particles/Particles_sinkSortParticles
Driver/Driver_finalizeFlash	Particles/Particles_sinkSumAttributes
Driver/Driver_finalizeSourceTerms	Particles/Particles_sinkSyncWithParticles
Driver/Driver_getComm	Particles/Particles_specifyMethods
Driver/Driver_getDt	Particles/Particles_unitTest
Driver/Driver_getElapsedWCTime	Particles/Particles_updateAttributes
Driver/Driver_getMType	Particles/Particles_updateGridVar
Driver/Driver_getNStep	Particles/Particles_updateRefinement
Driver/Driver_getNumProcs	PhysicalConstants/PhysicalConstants_get
Driver/Driver_getSimTime	PhysicalConstants/PhysicalConstants_init
Driver/Driver_getTimeStamp	PhysicalConstants/PhysicalConstants_list
Driver/Driver_init	PhysicalConstants/PhysicalConstants_listUnits
Driver/Driver_initFlash	PhysicalConstants/PhysicalConstants_unitTest
Driver/Driver_initMaterialProperties	physics/Cosmology/Cosmology_cdmPowerSpectrum
Driver/Driver_initNumericalTools	physics/Cosmology/Cosmology_computeDeltaCrit
Driver/Driver_initParallel	physics/Cosmology/Cosmology_computeDt
Driver/Driver_initSourceTerms	physics/Cosmology/Cosmology_computeDt
Driver/Driver_logMemoryUsage	physics/Cosmology/Cosmology_computeVariance
Driver/Driver_mpiThreadSupport	physics/Cosmology/Cosmology_finalize
Driver/Driver_putTimeStamp	physics/Cosmology/Cosmology_getOldRedshift
Driver/Driver_sendOutputData	physics/Cosmology/Cosmology_getParams
Driver/Driver_setupParallelEnv	physics/Cosmology/Cosmology_getRedshift
Driver/Driver_sourceTerms	physics/Cosmology/Cosmology_init
Driver/Driver_superTimeStep	physics/Cosmology/Cosmology_massToLength
Driver/Driver_verifyInitDt	physics/Cosmology/Cosmology_redshiftHydro
flashUtilities/Pipeline/Pipeline_finalize	physics/Cosmology/Cosmology_redshiftToTime
flashUtilities/Pipeline/Pipeline_globalCheckStatus	physics/Cosmology/Cosmology_sendOutputData
flashUtilities/Pipeline/Pipeline_globalCheckStructure	physics/Cosmology/Cosmology_solveFriedmannEqn
flashUtilities/Pipeline/Pipeline_init	physics/Cosmology/Cosmology_unitTest
flashUtilities/Pipeline/Pipeline_localActivate	physics/Diffuse/Diffuse
flashUtilities/Pipeline/Pipeline_localCreate	physics/Diffuse/Diffuse_advanceID
flashUtilities/Pipeline/Pipeline_localDeactivate	physics/Diffuse/Diffuse_computeDt
	physics/Diffuse/Diffuse_computeFluxLimiter

Happy coding!